

# Algorithmic needs for Fusion Magnetohydrodynamics (MHD) and other Predominantly Hyperbolic Systems of Equations

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## Outline

- Fusion...the ultimate energy source
- Basics of Magnetic Fusion
- The Role of Computer Simulation
- Implicit Solution of Hyperbolic Equations
- Approach of the CEMM codes
- A proposed 3D Solution Technique
- Other efforts within CEMM

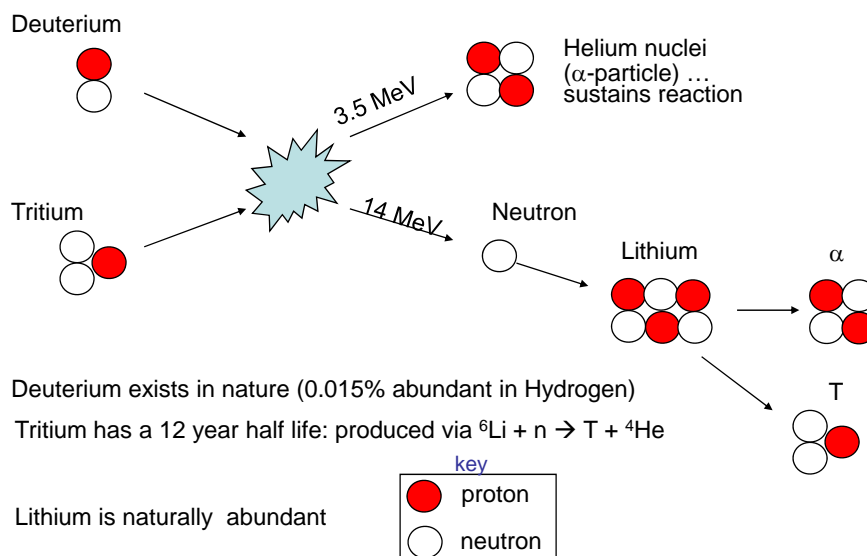


## **The Case for Fusion Energy**

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- Worldwide demand for energy continues to increase
  - Due to population increases and economic development
  - Most population growth and energy demand is in urban areas
    - Implies need for large, centralized power generation
- Worldwide oil and gas production is near or past peak
  - Need for alternative source: coal, fission, fusion
- Increasing evidence that release of greenhouse gases is causing global climate change . . . “Global warming”
  - Historical data and 100+ year detailed climate projections
  - This makes nuclear (fission or fusion) preferable to fossil (coal)
- Fusion has some advantages over fission that could become critical:
  - Inherent safety (no China syndrome)
  - No weapons proliferation considerations (security)
  - Greatly reduced waste disposal problems (no Yucca Mt.)

## Controlled Fusion uses isotopes of Hydrogen in a High Temperature Ionized Gas (Plasma)



## Controlled Fusion Basics

Create a mixture of D and T (plasma), heat it to high temperature, and the D and T will fuse to produce energy.

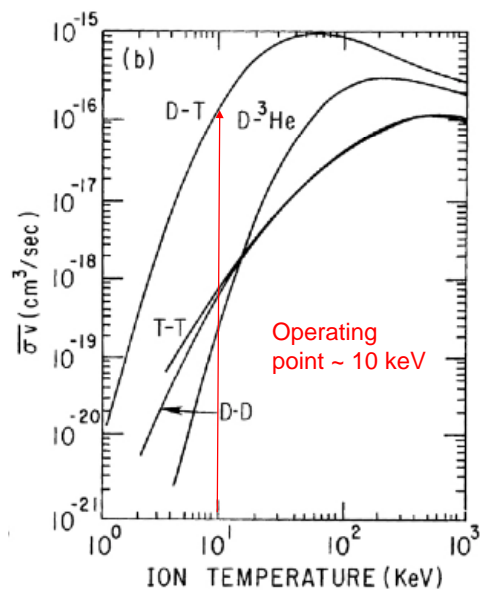
$$P_{DT} = n_D n_T \langle \sigma v \rangle (U_\alpha + U_n)$$

at 10 keV,  $\langle \sigma v \rangle \sim T^2$

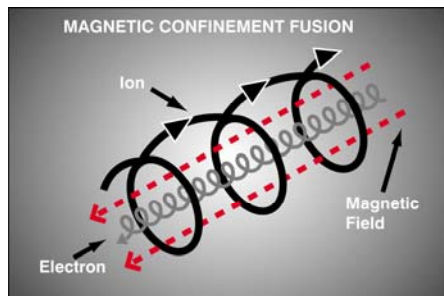
$$P_{DT} \sim (\text{plasma pressure})^2$$

**Need ~ 5 atmosphere @ 10 keV**

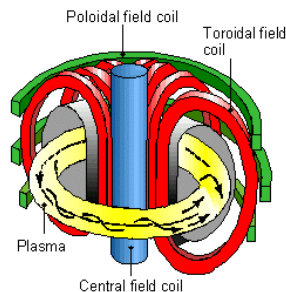
Note: 1 keV = 10,000,000 deg(K)



## Toroidal Magnetic Confinement



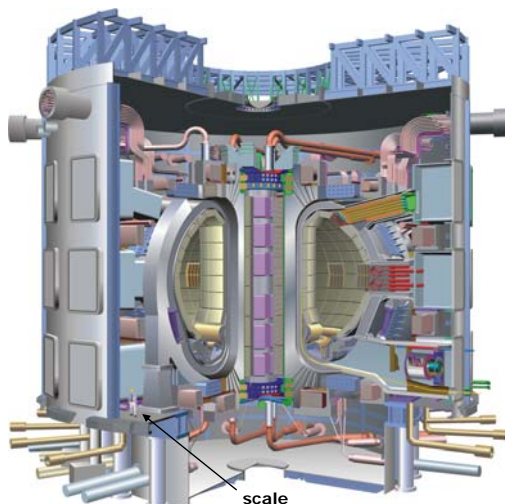
Charged particles have helical orbits in a magnetic field; they describe circular orbits perpendicular to the field and free-stream in the direction of the field.



TOKAMAK creates toroidal magnetic fields to confine particles in the 3<sup>rd</sup> dimension. Includes an induced toroidal plasma current to heat and confine the plasma

“TOKAMAK”: Russian abbreviation for “toroidal chamber”

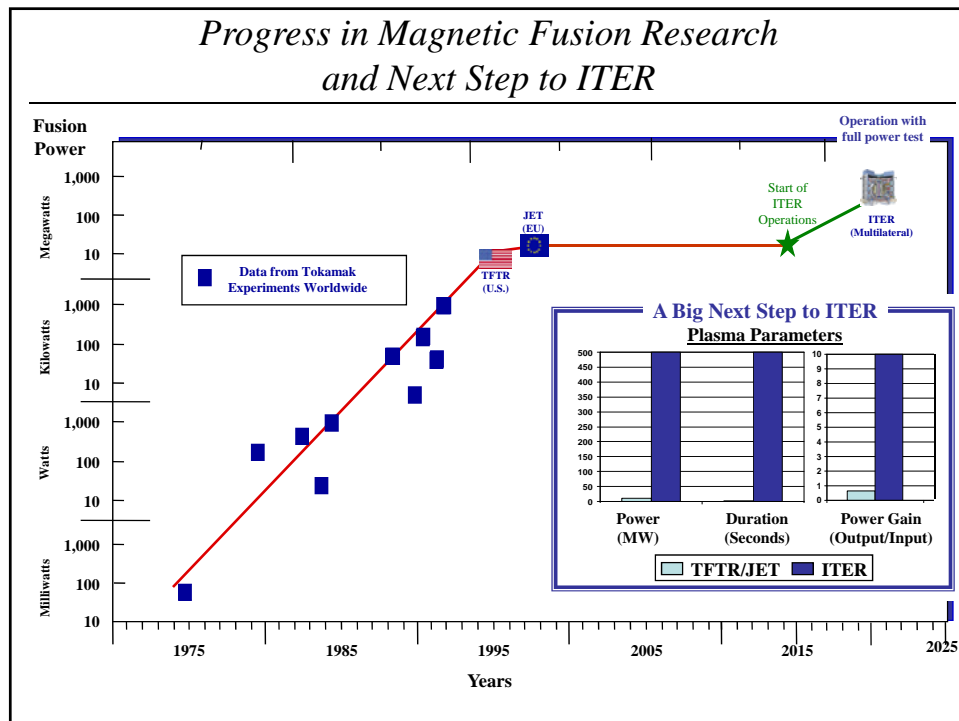
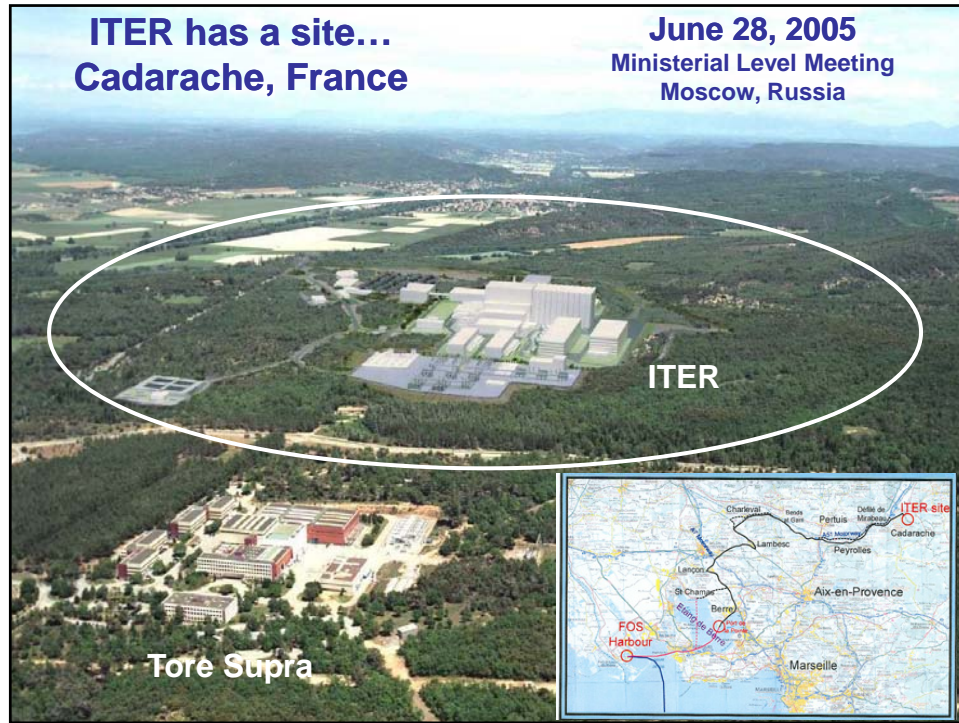
## The U.S. is an official partner in ITER



- 500 MW fusion output
- Cost: \$ 5-10 B
- To begin operation in 2015

International  
Thermonuclear  
Experimental Reactor:

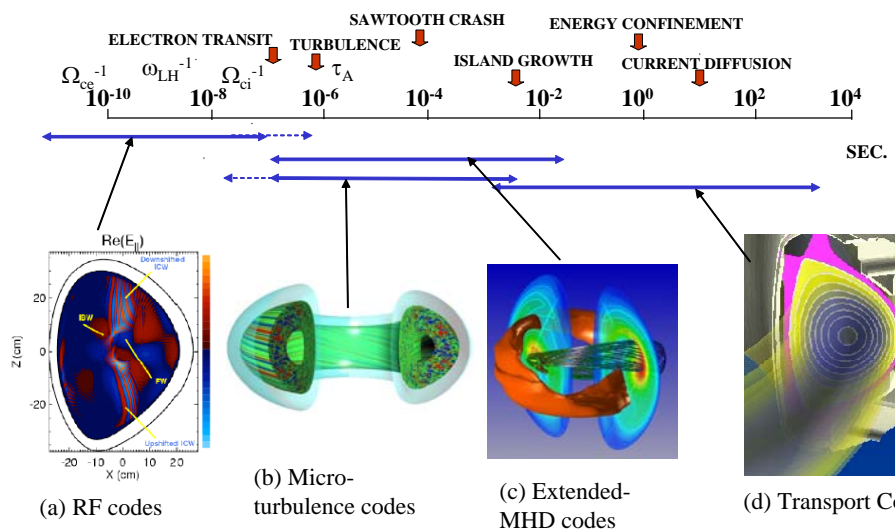
- European Union
- Japan
- United States
- Russia
- Korea
- China
- India
- World's largest tokamak
- all super-conducting coils



## Simulations are needed in 4 areas

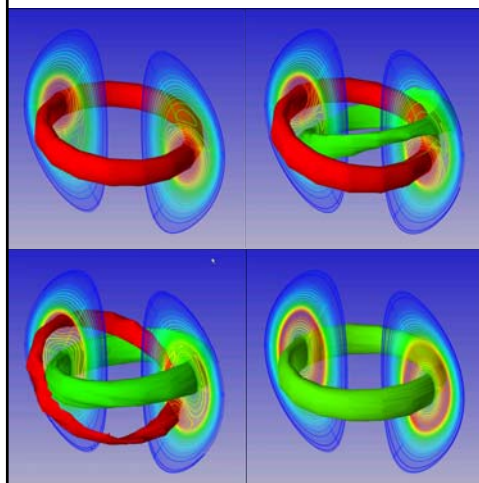
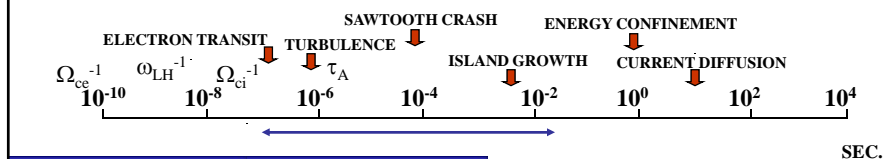
- How to heat the plasma to thermonuclear temperatures (  $\sim 100,000,000^\circ\text{C}$  )
- How to reduce the background turbulence
- How to eliminate device-scale instabilities
- How to optimize the operation of the whole device

These 4 areas address different timescales and are normally studied using different codes



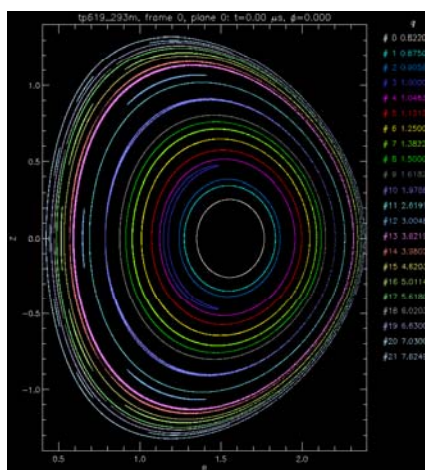


### Extended MHD Codes solve 3D fluid equations for device-scale stability



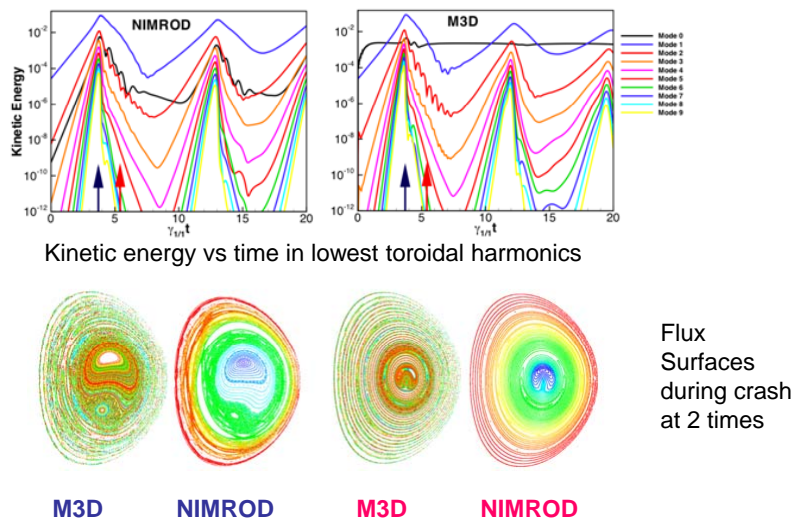
- Sawtooth cycle is one example of global phenomena that need to be understood
- Can cause degradation of confinement, or plasma termination
- The X-MHD codes typically exhibit good parallel scaling to 500-1000 processors
- Running time is dominated by elliptic solves
- Need to run for many time steps

## Show Quicktime Movie



- Example of a recent 3D CEMM calculation using M3D code
- "Internal Kink" mode in a small tokamak (Sawtooth Oscillations)
- Good agreement between M3D, NIMROD, and experimental results
- 500 wallclock hours and over 200,000 CPU-hours

## Excellent Agreement between NIMROD and M3D



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Consider a simple 1-D Hyperbolic System of Equations (Wave Equation)

$$\left. \begin{aligned} \frac{\partial u}{\partial t} &= c \frac{\partial v}{\partial x} \\ \frac{\partial v}{\partial t} &= c \frac{\partial u}{\partial x} \end{aligned} \right\} \frac{\partial^2 u}{\partial t^2} - c^2 \frac{\partial^2 u}{\partial x^2} = 0$$

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Implicit Centered Difference:

$$\frac{u_j^{n+1} - u_j^n}{\delta t} = c \left[ \theta \left( \frac{v_{j+1/2}^{n+1} - v_{j-1/2}^{n+1}}{\delta x} \right) + (1-\theta) \left( \frac{v_{j+1/2}^n - v_{j-1/2}^n}{\delta x} \right) \right]$$

$$\frac{v_{j+1/2}^{n+1} - v_{j+1/2}^n}{\delta t} = c \left[ \theta \left( \frac{u_{j+1}^{n+1} - u_j^{n+1}}{\delta x} \right) + (1-\theta) \left( \frac{u_{j+1}^n - u_j^n}{\delta x} \right) \right]$$

Stable for  $\theta \geq \frac{1}{2}$  2<sup>nd</sup> order for  $\theta = \frac{1}{2}$

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$$\frac{v_{j+1/2}^{n+1} - v_{j+1/2}^n}{\delta t} = c \left[ \theta \left( \frac{u_{j+1}^{n+1} - u_j^{n+1}}{\delta x} \right) + (1-\theta) \left( \frac{u_{j+1}^n - u_j^n}{\delta x} \right) \right]$$

Substitute from second equation into first:

$$u_j^{n+1} = u_j^n + (\delta t c)^2 \left[ \theta^2 \left( \frac{u_{j+1}^{n+1} - 2u_j^{n+1} + u_{j-1}^{n+1}}{\delta x^2} \right) + \theta(1-\theta) \left( \frac{u_{j+1}^n - 2u_j^n + u_{j-1}^n}{\delta x^2} \right) \right] + \delta t c \left( \frac{v_{j+1/2}^n - v_{j-1/2}^n}{\delta x} \right)$$

$$v_{j+1/2}^{n+1} = v_{j+1/2}^n + \frac{\delta t c}{\delta x} \left[ \theta (u_{j+1}^{n+1} - u_j^{n+1}) + (1-\theta) (u_{j+1}^n - u_j^n) \right]$$

These two equations can be solved sequentially

Only first involves Matrix Inversion ... Diagonally Dominant

$$u_j^{n+1} = u_j^n + (\delta t c)^2 \left[ \theta^2 \left( \frac{u_{j+1}^{n+1} - 2u_j^{n+1} + u_{j-1}^{n+1}}{\delta x^2} \right) + \theta(1-\theta) \left( \frac{u_{j+1}^n - 2u_j^n + u_{j-1}^n}{\delta x^2} \right) \right] + \delta t c \left( \frac{v_{j+1/2}^n - v_{j-1/2}^n}{\delta x} \right)$$

$$v_{j+1/2}^{n+1} = v_{j+1/2}^n + \frac{\delta t c}{\delta x} \left[ \theta (u_{j+1}^{n+1} - u_j^{n+1}) + (1-\theta) (u_{j+1}^n - u_j^n) \right]$$

Rewrite using standard finite-difference notation:

$$\delta_x^2 u_j^{n+1} \equiv u_{j+1}^{n+1} - 2u_j^{n+1} + u_{j-1}^{n+1}$$

$$\delta_x v_j^n \equiv v_{j+1/2}^n - v_{j-1/2}^n$$

$$S \equiv \frac{\delta t c}{\delta x}$$

Operator to invert

$$\boxed{1 - S^2 \theta^2 \delta_x^2} u_j^{n+1} = \left[ 1 + S^2 \theta(1-\theta) \delta_x^2 \right] u_j^n + S \delta_x v_j^n$$

$$v_{j+1/2}^{n+1} = v_{j+1/2}^n + S \left[ \theta \delta_x u_{j+1/2}^{n+1} + (1-\theta) \delta_x u_{j+1/2}^n \right]$$

This derivation can also be done in block matrix form:

Original equation:

$$\begin{bmatrix} \mathbf{I} & \mathbf{A}_{12} \\ \mathbf{A}_{21} & \mathbf{I} \end{bmatrix} \begin{bmatrix} \mathbf{u} \\ \mathbf{v} \end{bmatrix}^{n+1} = \begin{bmatrix} \mathbf{B}_{11} & \mathbf{B}_{12} \\ \mathbf{B}_{21} & \mathbf{B}_{22} \end{bmatrix} \begin{bmatrix} \mathbf{u} \\ \mathbf{v} \end{bmatrix}^n$$

$$\mathbf{A}_{12} = \begin{bmatrix} x & & & \\ x & x & & \\ & x & x & \\ & & x & x \end{bmatrix} \quad \mathbf{A}_{21} = \begin{bmatrix} x & x & & \\ & x & x & \\ & & x & x \\ & & & x \end{bmatrix}$$

Lower and upper bi-diagonal matrices

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Lower and upper bi-diagonal matrices

Shur complement:

$$\mathbf{v}^{n+1} = -\mathbf{A}_{21}\mathbf{u}^{n+1} + \mathbf{B}_{21}\mathbf{u}^n + \mathbf{B}_{22}\mathbf{v}^n$$

$$\mathbf{u}^{n+1} = \mathbf{A}_{12}\mathbf{A}_{21}\mathbf{u}^{n+1} + (\mathbf{B}_{11} - \mathbf{A}_{12}\mathbf{B}_{21})\mathbf{u}^n + (\mathbf{B}_{12} - \mathbf{A}_{12}\mathbf{B}_{22})\mathbf{v}^n$$

or:

$$\boxed{\mathbf{I} - \mathbf{A}_{12}\mathbf{A}_{21}} \mathbf{u}^{n+1} = (\mathbf{B}_{11} - \mathbf{A}_{12}\mathbf{B}_{21})\mathbf{u}^n + (\mathbf{B}_{12} - \mathbf{A}_{12}\mathbf{B}_{22})\mathbf{v}^n$$

$$\mathbf{v}^{n+1} = -\mathbf{A}_{21}\mathbf{u}^{n+1} + \mathbf{B}_{21}\mathbf{u}^n + \mathbf{B}_{22}\mathbf{v}^n$$

Operator to invert (tridiagonal for 1D)

An alternate derivation:

$$\begin{array}{ccc} \frac{\partial u}{\partial t} = c \frac{\partial v}{\partial x} & \longrightarrow & \frac{\partial u}{\partial t} = c \frac{\partial}{\partial x} \left[ v^n + \theta \delta t \frac{\partial v}{\partial t} \right] \\ \frac{\partial v}{\partial t} = c \frac{\partial u}{\partial x} & \text{Expand RHS in Taylor} & \frac{\partial v}{\partial t} = c \frac{\partial}{\partial x} \left[ u^n + \theta \delta t \frac{\partial u}{\partial t} \right] \\ & \text{series in time to time-} & \\ & \text{center} & \end{array}$$

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$$\begin{array}{l} \frac{\partial u}{\partial t} = c \frac{\partial}{\partial x} \left[ v^n + \theta \delta t \left( c \frac{\partial}{\partial x} \left[ u^n + \theta \delta t \frac{\partial u}{\partial t} \right] \right) \right] \\ \frac{\partial v}{\partial t} = c \frac{\partial}{\partial x} \left[ u^n + \theta \delta t \frac{\partial u}{\partial t} \right] \end{array}$$

Substitute from second equation into first

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Expand RHS in Taylor series in time to time-center

$$\begin{aligned} \frac{\partial u}{\partial t} &= c \frac{\partial}{\partial x} \left[ v^n + \theta \delta t \left( c \frac{\partial}{\partial x} \left[ u^n + \theta \delta t \frac{\partial u}{\partial t} \right] \right) \right] \\ \frac{\partial v}{\partial t} &= c \frac{\partial}{\partial x} \left[ u^n + \theta \delta t \frac{\partial u}{\partial t} \right] \end{aligned}$$

Substitute from second equation into first

Use standard centered difference in time:

$$\begin{aligned} \left[ 1 - \theta^2 (\delta t)^2 c^2 \frac{\partial^2}{\partial x^2} \right] u^{n+1} &= \left[ 1 + \theta(1-\theta)(\delta t)^2 c^2 \frac{\partial^2}{\partial x^2} \right] u^n + \delta t c \frac{\partial}{\partial x} v^n \\ v^{n+1} &= v^n + \delta t c \left[ \theta \frac{\partial}{\partial x} u^{n+1} + (1-\theta) \frac{\partial}{\partial x} u^n \right] \end{aligned}$$

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This is the same operator as before when centered spatial differences used

Now apply this technique to the basic 3D MHD equations:

$$\left. \begin{aligned} \rho_0 \dot{\mathbf{V}} &= \frac{1}{\mu_0} [\nabla \times \mathbf{B}] \times \mathbf{B} - \nabla p \\ \dot{\mathbf{B}} &= \nabla \times [\mathbf{V} \times \mathbf{B}] \\ \dot{p} &= -\mathbf{V} \cdot \nabla p - \gamma p \nabla \cdot \mathbf{V} \end{aligned} \right\} \begin{array}{l} \text{Ideal MHD Equations for velocity,} \\ \text{magnetic field, and pressure:} \\ \text{Symmetric Hyperbolic System} \\ \text{7-waves} \end{array}$$

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$$\begin{aligned} \rho_0 \dot{\mathbf{V}} &= \frac{1}{\mu_0} [\nabla \times (\mathbf{B} + \theta \delta t \dot{\mathbf{B}})] \times (\mathbf{B} + \theta \delta t \dot{\mathbf{B}}) - \nabla (p + \theta \delta t \dot{p}) \\ \dot{\mathbf{B}} &= \nabla \times [(\mathbf{V} + \theta \delta t \dot{\mathbf{V}}) \times \mathbf{B}] \\ \dot{p} &= -(\mathbf{V} + \theta \delta t \dot{\mathbf{V}}) \cdot \nabla p - \gamma p \nabla \cdot (\mathbf{V} + \theta \delta t \dot{\mathbf{V}}) \end{aligned}$$

Taylor Expand in  
Time as before



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Taylor Expand in Time as before

Substitute from 2<sup>nd</sup> and 3<sup>rd</sup> equation into first, finite difference in time:

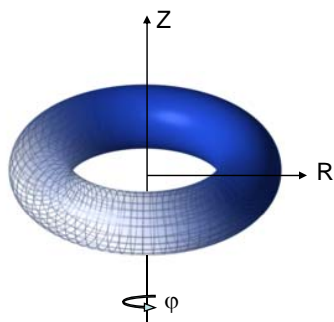
$$\{\rho - \theta^2 (\delta t)^2 L\} \mathbf{V}^{n+1} = \{\rho - \theta(\theta-1)(\delta t)^2 L\} \mathbf{V}^n + \delta t \left\{ -\nabla p + \frac{1}{\mu_0} (\nabla \times \mathbf{B}) \times \mathbf{B} \right\}$$

MHD Operator:  $\longrightarrow$

$$L\{\mathbf{V}\} = \frac{1}{\mu_0} \left\{ \nabla \times [\nabla \times (\mathbf{V} \times \mathbf{B})] \right\} \times \mathbf{B} + \frac{1}{\mu_0} (\nabla \times \mathbf{B}) \times [\nabla \times (\mathbf{V} \times \mathbf{B})] + \nabla (\mathbf{V} \cdot \nabla p + \gamma p \nabla \cdot \mathbf{V})$$

$$\{\rho - \theta^2 (\delta t)^2 L\} \mathbf{V}^{n+1} = \{\rho - \theta(\theta-1)(\delta t)^2 L\} \mathbf{V}^n + \delta t \left\{ -\nabla p + \frac{1}{\mu_0} (\nabla \times \mathbf{B}) \times \mathbf{B} \right\}$$

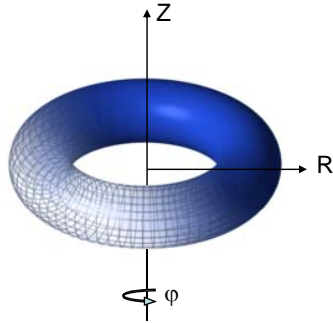
$$L\{\mathbf{V}\} = \frac{1}{\mu_0} \left\{ \nabla \times [\nabla \times (\mathbf{V} \times \mathbf{B})] \right\} \times \mathbf{B} + \frac{1}{\mu_0} (\nabla \times \mathbf{B}) \times [\nabla \times (\mathbf{V} \times \mathbf{B})] + \nabla (\mathbf{V} \cdot \nabla p + \gamma p \nabla \cdot \mathbf{V})$$



- Need to solve this in 3D torus with strong magnetic field in toroidal direction ( $\phi$ ) ... anisotropy
- Wide range of wave speeds leads to ill-conditioned matrices
- Gradients in (R,Z) plane much larger than in  $\phi$  direction
- Also need to preserve  $\nabla \cdot \mathbf{B} = 0$

$$\{\rho - \theta^2 (\delta t)^2 L\} \mathbf{V}^{n+1} = \{\rho - \theta(\theta-1)(\delta t)^2 L\} \mathbf{V}^n + \delta t \left\{ -\nabla p + \frac{1}{\mu_0} (\nabla \times \mathbf{B}) \times \mathbf{B} \right\}$$

$$L\{\mathbf{V}\} = \frac{1}{\mu_0} \left\{ \nabla \times [\nabla \times (\mathbf{V} \times \mathbf{B})] \right\} \times \mathbf{B} + \frac{1}{\mu_0} (\nabla \times \mathbf{B}) \times [\nabla \times (\mathbf{V} \times \mathbf{B})] \\ + \nabla (\mathbf{V} \cdot \nabla p + \gamma p \nabla \cdot \mathbf{V})$$



CEMM has 3 Major Codes using different approaches for solving this equation:

**NIMROD:** Fourier analyses in  $\phi$  and solves each harmonic separately using SuperLU with GMRES to couple harmonics

**M3D:** Finite Differences in  $\phi$ , but only solves implicit operator equation in (R,Z) planes, using either GMRES or HYPRE (via PetSC) Explicit differences in  $\phi$ .

**M3D-C':** Finite Differences in  $\phi$ , uses SuperLU in (R,Z) plane, now going to 3D.

M3D-C' uses a stream function / potential form of the velocity and magnetic fields

$$\vec{V} = \nabla U \times \hat{z} + \nabla_{\perp} \chi + V_z \hat{z}$$

$$\vec{B} = \nabla \psi \times \hat{z} + I \hat{z}$$

The sparse matrix equation to be solved for the velocity variables take the form:

$$\begin{bmatrix} S_{11}^v & S_{12}^v & S_{13}^v \\ S_{21}^v & S_{22}^v & S_{23}^v \\ S_{31}^v & S_{32}^v & S_{33}^v \end{bmatrix} \begin{bmatrix} U \\ V_z \\ \chi \end{bmatrix}^{n+1} = \begin{bmatrix} D_{11}^v & D_{12}^v & D_{13}^v \\ D_{21}^v & D_{22}^v & D_{23}^v \\ D_{31}^v & D_{32}^v & D_{33}^v \end{bmatrix} \begin{bmatrix} U \\ V_z \\ \chi \end{bmatrix}^n + \begin{bmatrix} R_{11}^v & R_{12}^v & R_{13}^v \\ R_{21}^v & R_{22}^v & R_{23}^v \\ R_{31}^v & R_{32}^v & R_{33}^v \end{bmatrix} \begin{bmatrix} \psi \\ I \\ p_e \end{bmatrix}^n$$

- Corresponds to the operator equation derived on earlier vg:

$$\{\rho - \theta^2 (\delta t)^2 L\} \mathbf{V}^{n+1} = \{\rho - \theta(\theta-1)(\delta t)^2 L\} \mathbf{V}^n + \delta t \left\{ -\nabla p + \frac{1}{\mu_0} (\nabla \times \mathbf{B}) \times \mathbf{B} \right\}$$

- Also contains 2 non-trivial sub-systems (reduced MHD)

$$\begin{bmatrix} S_{11}^v \end{bmatrix} [U]^{n+1} = \begin{bmatrix} D_{11}^v \end{bmatrix} [U]^n + \begin{bmatrix} R_{11}^v \end{bmatrix} [\psi]^n \quad \text{etc.}$$

## In 2D, Implicit equations are solved using SuperLU\_Dist

Details for bassi.nersc.gov:

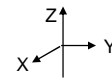
Mesh points	180 x 180	# processors	8	32	128
Matrix Rank	$5.9 \times 10^5$	Factor (s)	69.5	38.1	16.9
# Non-zeros	$9.5 \times 10^7$	Gflop/s	27.2	50.1	112.8*
# NZ in L/U	$8.8 \times 10^8$				

Total problem time (8 processors) for typical high resolution reconnection problem = 208 s x 400 cycles x 8p = 185 p-hrs

Note that for **linear problem**, Matrix need only be factored once. For **semi-implicit** method, matrix needs to be factored only occasionally.

\*NOTE: In 3D, if we had 100 planes with simultaneous instances of SuperLU, this would be 12,800 p and 11.2 Tflop/s actual!

## How to take this to 3D ?



Consider (1,1) component in 3D: Corresponds to "Strauss Equations" (3D reduced MHD for  $\psi$  and  $U$ )

note: "Poisson Bracket"

$$[a,b] \equiv \frac{\partial a}{\partial X} \frac{\partial b}{\partial Y} - \frac{\partial a}{\partial Y} \frac{\partial b}{\partial X}$$

"Inner Product"

$$(a,b) \equiv \frac{\partial a}{\partial X} \frac{\partial b}{\partial X} + \frac{\partial a}{\partial Y} \frac{\partial b}{\partial Y}$$

$$\nabla_{\perp}^2 \dot{U} + [\nabla_{\perp}^2 U, U] = [\nabla_{\perp}^2 \psi, \psi] + B \frac{\partial}{\partial z} \nabla_{\perp}^2 \psi$$

$$\dot{\psi} + [\psi, U] = B \frac{\partial}{\partial z} U$$

Taylor expand in time:

$$\begin{aligned} \nabla_{\perp}^2 \dot{U} + [\nabla_{\perp}^2 U + \theta \delta t \nabla_{\perp}^2 \dot{U}, U + \theta \delta t \dot{U}] &= [\nabla_{\perp}^2 \psi + \theta \delta t \nabla_{\perp}^2 \dot{\psi}, \psi + \theta \delta t \dot{\psi}] \\ &\quad + B \frac{\partial}{\partial z} \nabla_{\perp}^2 \psi + \theta \delta t B \frac{\partial}{\partial z} \nabla_{\perp}^2 \dot{\psi} \end{aligned}$$

$$\dot{\psi} + [\psi, U + \theta \delta t \dot{U}] = B \frac{\partial}{\partial z} U + \theta \delta t B \frac{\partial}{\partial z} \dot{U}$$

Substitute field derivatives into velocity equation to get implicit equation:

Finite Difference in Z:  $\dot{U}_z = \frac{1}{\delta z} [\dot{U}_{j+1} - \dot{U}_{j-1}]$   $\dot{U}_{zz} = \frac{1}{(\delta z)^2} [\dot{U}_{j+1} - 2\dot{U}_j + \dot{U}_{j-1}]$

When we apply  $C^r$  continuous finite elements in  $(x, y)$ , we get a block tridiagonal equation, with the matrix blocks being 2D matrices:

$$B_j^0 U_j^{n+1} + D_j^0 + \varepsilon (A_j^1 U_{j+1}^{n+1} + B_j^1 U_j^{n+1} + C_j^1 U_{j-1}^{n+1} + D_j^1) = 0$$

$$B_j^0 U_j^{n+1} = v_i \nabla_\perp^2 \dot{U}_j - (\theta \delta t)^2 \left[ + \nabla_\perp^2 \psi [v_i, [\psi, U_j^{n+1}]] - ([\psi, U_j^{n+1}], [v_i, \psi]) + B(v_i, [\psi_z, U_j^{n+1}]) \right]$$

$$B_j^1 U_j^{n+1} = v_i \nabla_\perp^2 \dot{U}_j - (\theta \delta t)^2 \left[ - \frac{2B^2}{(\delta z)^2} v_i \nabla_\perp^2 U_j^{n+1} \right]$$

$$A_j U_{j+1}^{n+1} = -(\theta \delta t)^2 \left[ - \frac{B}{\delta z} \nabla_\perp^2 \psi [v_i, U_{j+1}^{n+1}] - \frac{B}{\delta z} \nabla_\perp^2 U_{j+1}^{n+1} [v_i, \psi] + \frac{B^2}{(\delta z)^2} v_i \nabla_\perp^2 U_{j+1}^{n+1} + \frac{B}{\delta z} (v_i, [\psi, U_{j+1}^{n+1}]) \right]$$

$$C_j U_{j-1}^{n+1} = -(\theta \delta t)^2 \left[ \frac{B}{\delta z} \nabla_\perp^2 \psi [v_i, U_{j-1}^{n+1}] + \frac{B}{\delta z} \nabla_\perp^2 U_{j-1}^{n+1} [v_i, \psi] + \frac{B^2}{(\delta z)^2} v_i \nabla_\perp^2 U_{j-1}^{n+1} - \frac{B}{\delta z} (v_i, [\psi, U_{j-1}^{n+1}]) \right]$$

Can this structure be used to define an efficient iteration scheme where the 2D direct solves serve as a preconditioner ?

## Summary of Proposed 3D Time Advance

$$B_j^0 U_j^{n+1} + D_j^0 + \varepsilon (A_j^1 U_{j+1}^{n+1} + B_j^1 U_j^{n+1} + C_j^1 U_{j-1}^{n+1} + D_j^1) = 0$$

$U_j^{n+1}$  is vector of all unknown velocities on plane j at new time

$B_j^0, A_j^1, B_j^1, C_j^1$  are 2D sparse matrices at plane j

$D_j^0, D_j^1$  are 2D vectors at plane j

Possible iteration scheme. Use SuperLU to factor the  $B_j^0$  simultaneously

$$U_j^{i+1} = -[B_j^0]^{-1} [D_j^0 + \varepsilon (A_j^1 U_{j+1}^i + B_j^1 U_j^i + C_j^1 U_{j-1}^i + D_j^1)]$$

Note that  $B_j^0$  matrices only need to be factored once per timestep

## Other Efforts in CEMM

NIMROD:

Sovinec (U. Wisc):

- Now using “poloidal preconditioning”. Exploring ‘toroidal preconditioning

S. Vadlamani, S. Kruger (Tech X), T. Manteuffel, S. McCormick (CU APPM):

- SBIR contract to explore the use of MultiGrid on the MHD equations. Plan to use HYPRE via the PETSc interface